

Smooth Manifolds and Minkowski Spacetime

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3/2/11

Outline

- 1 Topology and Curves
- 2 Surfaces
- 3 Spaces and the Schwarzschild Metric

Topology and Curves

Definition

to•pol•o•gy |tə'päləjē|

noun

the study of geometric properties and spatial relations
unaffected by the continuous change of shape or size of figures.

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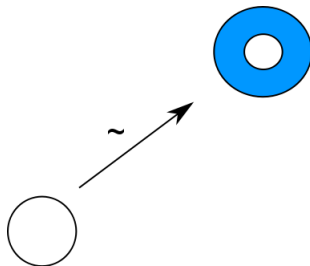
Remark

Actually, **topology** is the study of topological spaces (X, T) where X is a set and $T \supseteq \{\emptyset, X\}$ is a collection of subsets of X which is closed under arbitrary unions and finite intersections.

Types of Topological Equivalence

- Homotopy equivalent

circle \sim annulus



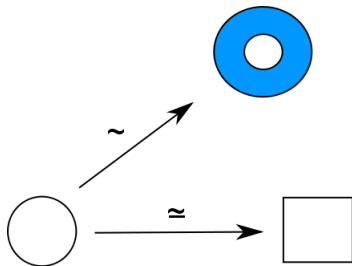
Types of Topological Equivalence

- Homotopy equivalent

circle \sim annulus

- Homeomorphic

circle \cong square



Types of Topological Equivalence

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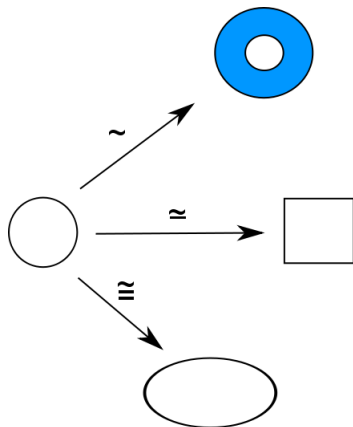
circle \sim annulus

- Homeomorphic

circle \cong square

- Diffeomorphic

circle \cong ellipse



Here we'll be interested in “manifolds” that are

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- connected (all one piece)

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- connected (all one piece)
- locally Euclidean (if we zoom in, things look flat)
- smooth (no sharp corners, so we can do calculus)

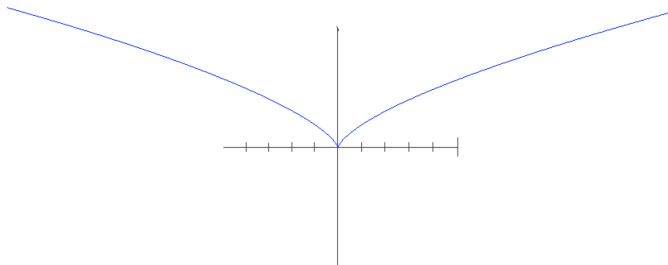


Figure: $y = x^{2/3}$ has a “cusp” at $(0,0)$

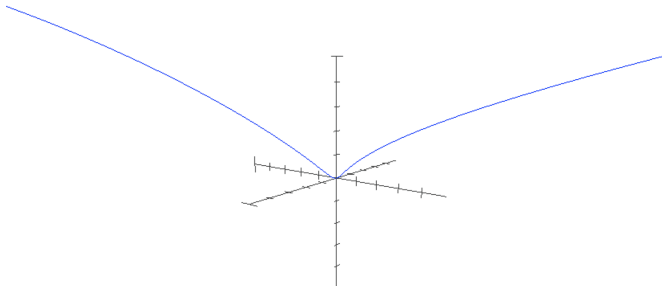
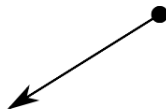


Figure: Maybe we're looking at a shadow of the twisted cubic

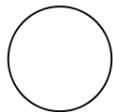
1-Dimensional Manifolds



line



ray



circle



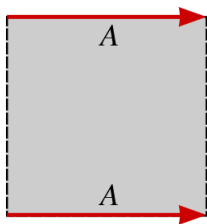
segment

Surfaces

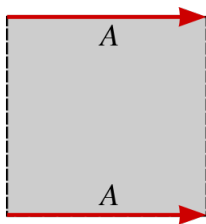
Polygon

Picture

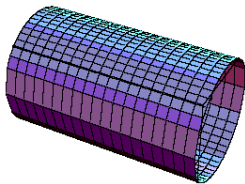
Boundary



Polygon



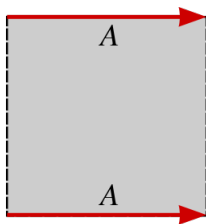
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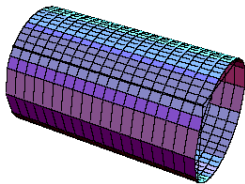
Boundary

Table: Finite Cylinder C

Polygon



Picture



Boundary

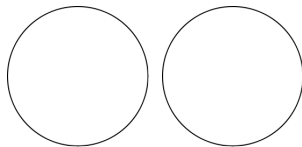
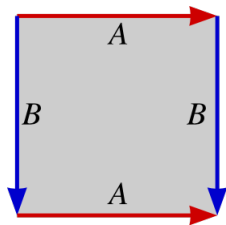


Table: Finite Cylinder C

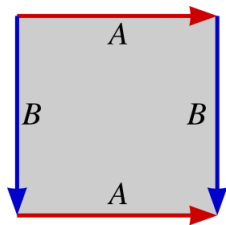
Polygon

Picture

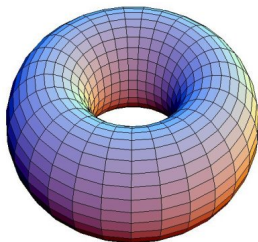
Boundary



Polygon



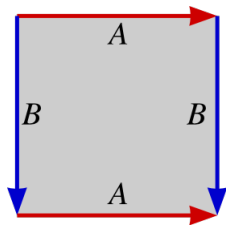
Picture



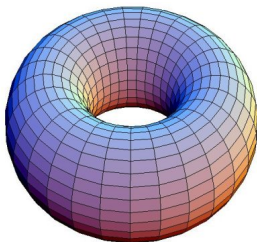
Boundary

Table: Torus \mathbb{T}^2

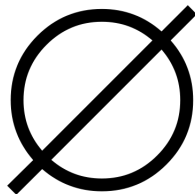
Polygon



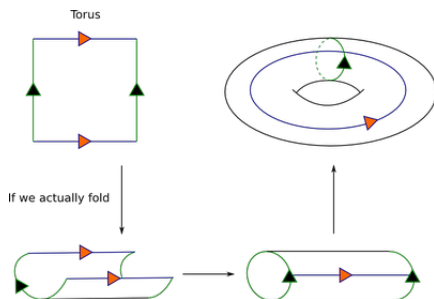
Picture



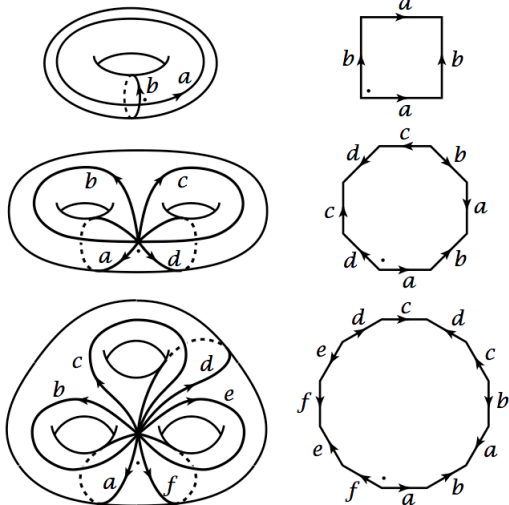
Boundary

Table: Torus \mathbb{T}^2

$$C \rightsquigarrow \mathbb{T}^2$$



Other Polygons Produce Multi-Holed Tori



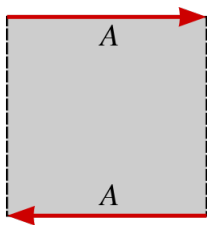
A Knotted Torus is Still a Torus



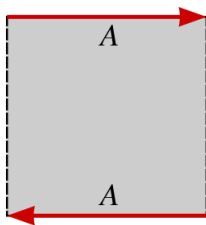
Polygon

Picture

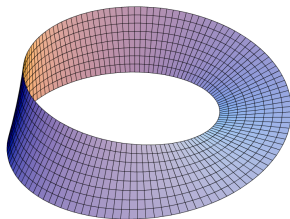
Boundary



Polygon



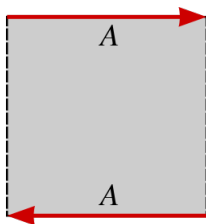
Picture



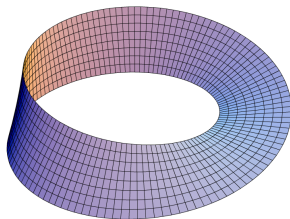
Boundary

Table: Möbius Strip M

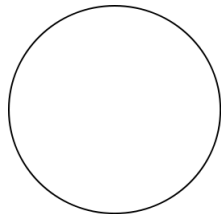
Polygon



Picture



Boundary

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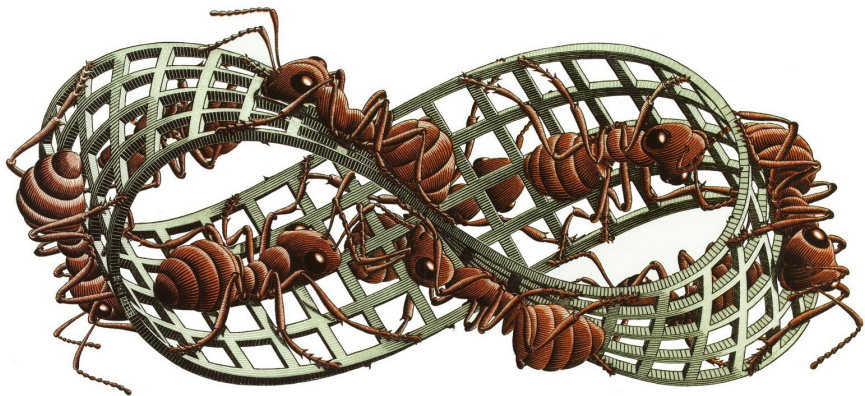
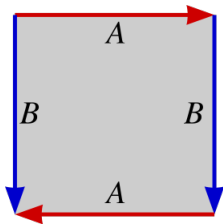


Figure: M. C. Escher, "Möbius Strip II"

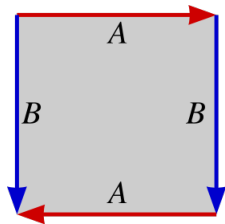
Polygon

Picture

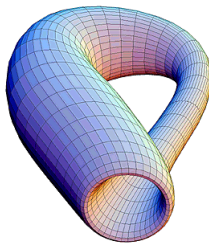
Boundary



Polygon



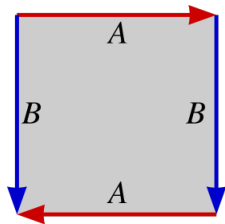
Picture



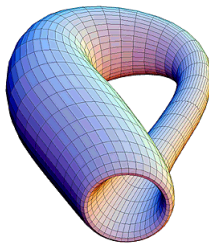
Boundary

Table: Klein Bottle K

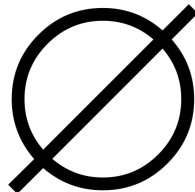
Polygon



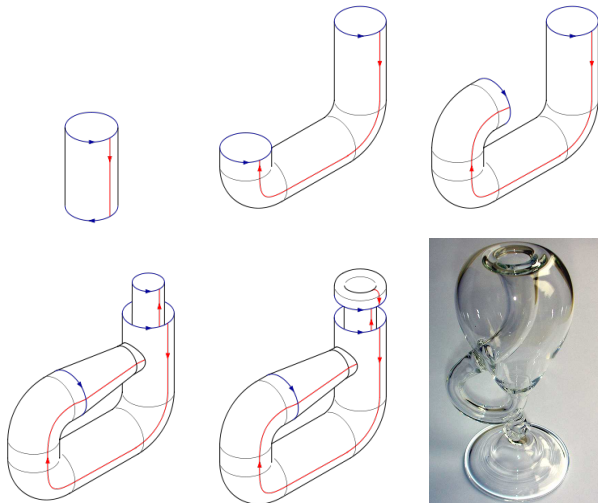
Picture



Boundary

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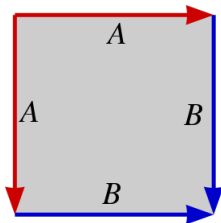
$$C \rightsquigarrow K$$



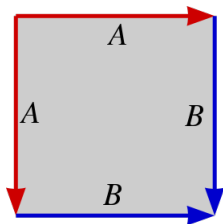
Polygon

Picture

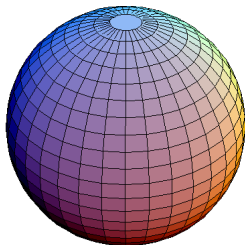
Boundary



Polygon



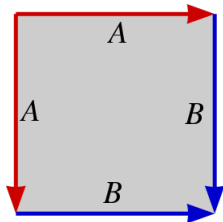
Picture



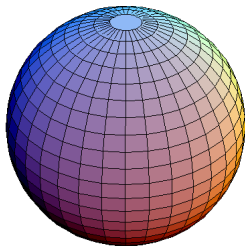
Boundary

Table: Sphere S^2

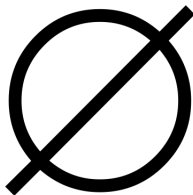
Polygon



Picture



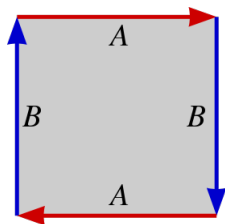
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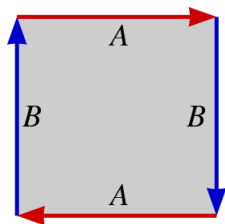
Polygon

Picture

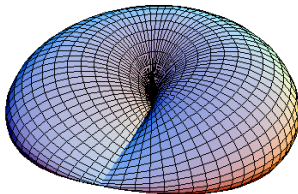
Boundary



Polygon



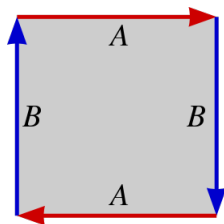
Picture



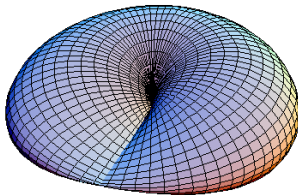
Boundary

Table: Real Projective Plane \mathbb{RP}^2

Polygon



Picture



Boundary

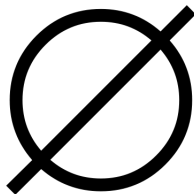
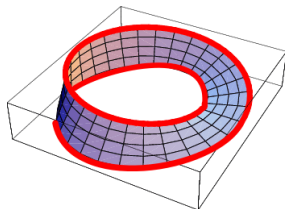
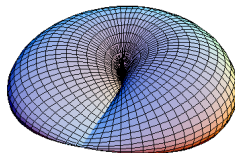
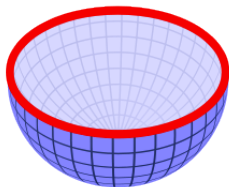


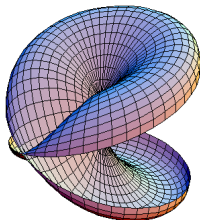
Table: Real Projective Plane \mathbb{RP}^2



↓ paste

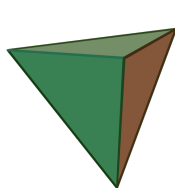


↓ slice

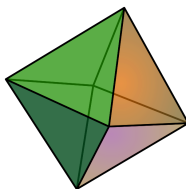


Spaces and the Schwarzschild Metric

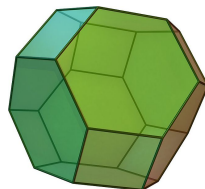
We can also take polyhedra (instead of polygons) and attach faces (instead of sides) to get 3-dimensional manifolds



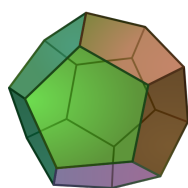
tetrahedron



octahedron



truncated oct.



dodecahedron

Table: Some Polyhedra

Minkowski Spacetime

We can indicate geometries with “line elements”

ds = an infinitesimal element of length

This allows us to do calculus: tangents, areas

Flat Spaces

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Flat Spaces

$$ds^2 = dx^2$$

linear

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Flat Spaces

$$ds^2 = dx^2$$

linear

$$ds^2 = dx^2 + dy^2$$

planar

Minkowski Spacetime

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Flat Spaces

$$ds^2 = dx^2$$

linear

$$ds^2 = dx^2 + dy^2$$

planar

$$ds^2 = dx^2 + dy^2 + dz^2$$

spatial

Minkowski Spacetime

We can indicate geometries with “line elements”

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This allows us to do calculus: tangents, areas

Flat Spaces

$$ds^2 = dx^2$$

linear

$$ds^2 = dx^2 + dy^2$$

planar

$$ds^2 = dx^2 + dy^2 + dz^2$$

spatial

$$ds^2 = dx^2 + dy^2 + dz^2 - dt^2$$

spatial-temporal

$$= -dt^2 + dr^2 + r^2(d\theta^2 + \sin^2(\theta)d\varphi^2)$$

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Einstein did this in 1915 via 10 “field equations.”

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Einstein did this in 1915 via 10 “field equations.”

In the case of a non-charged, non-rotating, uniformly dense, spherical body of mass M we get...

$$ds^2 = - \left(1 - \frac{2M}{r}\right) dt^2 + \left(1 - \frac{2M}{r}\right)^{-1} dr^2 + r^2(d\theta^2 + \sin^2(\theta)d\varphi^2)$$

Embedding diagrams help us visualize this geometry

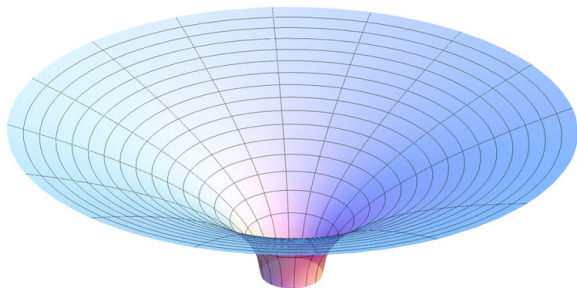


Figure: Flamm's Paraboloid

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It could reach an equilibrium supported by internal forces, or it may be unsupported against further gravitational collapse...

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If the radius shrinks below $2M$, a black hole is formed.

At a point in the life of a star, it runs out of thermonuclear fuel. It could reach an equilibrium supported by internal forces, or it may be unsupported against further gravitational collapse... If the radius shrinks below $2M$, a black hole is formed.



Figure: An artist's interpretation of a black hole